MARK SCHEME for the October/November 2010 question paper

for the guidance of teachers

9702 PHYSICS

9702/42 Paper 4 (A2 Structured Questions), maximum raw mark 100

This mark scheme is published as an aid to teachers and candidates, to indicate the requirements of the examination. It shows the basis on which Examiners were instructed to award marks. It does not indicate the details of the discussions that took place at an Examiners' meeting before marking began, which would have considered the acceptability of alternative answers.

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UNIVERSITY of CAMBRIDGE International Examinations

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			GCE AS	A LEVEL – Octo	bber/November	2010	9702	42		
	Section A									
1	(a) forc	ce per unit mass (<i>ratio idea essential</i>)							[1]	
	(b) gra		h: correct curvature from $(R, 1.0 g_S)$ & at least one other correct point							
	(c) (i)			•	posite directions			M1		
		<i>either</i> resultant field found by subtraction of the field strength or any other sensible comment so there is a point where it is zero (allow $F_E = -F_M$ for 2 marks)							[2]	
	(ii)) $GM_{\rm E} / x^2 = GM_{\rm M} / (D - x)^2$ (6.0 × 10 ²⁴) / (7.4 × 10 ²²) = x^2 / (60 $R_{\rm E} - x$) ² $x = 54 R_{\rm E}$								
	(iii)	(iii) graph: $g = 0$ at least $\frac{2}{3}$ distance to Moon g_E and g_M in opposite directions correct curvature (by eye) and $g_E > g_M$ at surface							[3]	
2	(a) (i)	no foi	rces (of attr	action or repulsic	on) between atom	ns / mole	cules / particle	s B1	[1]	
	(ii)		sum of kinetic and potential energy of atoms / molecules due to random motion						[2]	
	(iii)	(rand	om) kinetic	energy increases	s with temperatur	е		M1		
		no potential energy (so increase in temperature increases internal energy)							[2]	
	(b) (i)	zero						A1	[1]	
	(ii)	work	work done = $p\Delta V$							
			$= 4.0 \times 10^5 \times 6 \times 10^{-4}$ = 240 J (ignore any sign)							
	(iii)	(iii)								
	~ /	0	change work done / J heating / J increase in internal energy / J							
					000					

			energy / e
$\begin{array}{c} P \rightarrow Q \\ Q \rightarrow R \\ R \rightarrow P \end{array}$	+240 0 -840	-600 +720 +480	-360 +720 -360

(correct signs essential) (each horizontal line correct, 1 mark – max 3)

B3 [3]

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3	(a)	(i)	resor	nance		B1	[1]
		(ii)	ampl	itude 16mm and frequency 4.6Hz		A1	[1]
	(b)	(i)	a =	$(-)\omega^2 x \text{ and } \omega = 2\pi f$ $4\pi^2 \times 4.6^2 \times 16 \times 10^{-3}$ $13.4 \mathrm{m s^{-2}}$		C1 C1 A1	[3]
		(ii)	F =	<i>ma</i> 150 × 10 ^{−3} × 13.4		C1	
				2.0N		A1	[2]
	(c)			vs 'below' given line and never zero t 4.6 Hz (or slightly less) and flatter		M1 A1	[2]
4	(a)	cha	irge / j	potential (difference) (ratio must be clear)		B1	[1]
	(b)	(i)	V = 0	$Q / 4\pi \varepsilon_0 r$		B1	[1]
		(ii)	C = (so C	$Q/V = 4\pi \epsilon_0 r$ and $4\pi \epsilon_0$ is constant $\propto r$		M1 A0	[1]
	(c)	(i)	r = (6	/ $4\pi\epsilon_0 r$ 5.8×10^{-12}) / ($4\pi \times 8.85 \times 10^{-12}$) × 10^{-2} m		C1 C1 A1	[3]
		(ii)	Q = ($CV = 6.8 \times 10^{-12} \times 220$ = 1.5 × 10 ⁻⁹ C		A1	[1]
	(d)	(i)	V = (= 83	$Q/C = (1.5 \times 10^{-9}) / (18 \times 10^{-12})$		A1	[1]
		(ii)	eithe	$\Delta E = \frac{1}{2} \times 6.8 \times 10^{-12} \times 220^2 - \frac{1}{2} \times 18 \times 10^{-12} \times 83^2$		C1 C1	
			or	= $1.65 \times 10^{-7} - 6.2 \times 10^{-8}$ = 1.03×10^{-7} J energy = $\frac{1}{2}$ QV $\Delta E = \frac{1}{2} \times 1.5 \times 10^{-9} \times 220 - \frac{1}{2} \times 1.5 \times 10^{-9} \times 83$ = 1.03×10^{-7} J		A1 (C1) (C1) (A1)	[3]

	Pa			Mark Scheme: Teachers' version	Syllabus	Paper	
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5	(a)	field	eld into (the plane of) the paper			B1	[1]
	(b)		² / r = = (20	e to magnetic field <u>provides</u> the centripetal force Bqv 0 × 1.66 × 10 ⁻²⁷ × 1.40 × 10 ⁵) / (1.6 × 10 ⁻¹⁹ × 6.4 × 10 ⁻² 454 T	²)	B1 C1 B1 A0	[3]
	(c)	(i)	<u>sem</u>	icircle with diameter greater than 12.8 cm		B1	[1]
		(ii)	new	flux density = $\frac{22}{20} \times 0.454$		C1	
				B = 0.499 T		A1	[2]
6	(a)	(i)	e.g.	prevent flux losses / improve flux linkage		B1	[1]
		(ii)	e.m.	in core is changing f. / current (induced) <u>in core</u> ced current in core causes heating		B1 B1 B1	[3]
	(b)	(i)		value of the direct current producing same (mean) pov resistor	ver / heating	M1 A1	[2]
		(ii)	•	er in primary = power in secondary $P = V_S I_S$		M1 A1	[2]
7	(a)	(i)	e.g.	electron / particle diffraction		B1	[1]
		(ii)	e.g.	photoelectric effect		B1	[1]
	(b)	(i)	6			A1	[1]
		(ii)	λ = I = (6.	nge in energy = 4.57×10^{-19} J hc / E .63 × 10 ⁻³⁴ × 3.0 × 10 ⁸) / (4.57 × 10 ⁻¹⁹) 4 × 10 ⁻⁷ m		C1 A1	[2]
8	(0)					M1	[~]
ο	(a)	splitting of a heavy nucleus (<i>not atom/nuclide</i>) into two (lighter) nuclei of <u>approximately same mass</u>			A1	[2]	
	(b)	¹ n 42He 73Li		(allow $\frac{4}{2}\alpha$)		M2 A1	[3]
	(c)	emitted particles have kinetic energy		in rode /	B1		
		range of particles in the control rods is short / particles stopped in ro lose kinetic energy in rods kinetic energy of particles converted to thermal energy				B1 B1	[3]

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				Section B			
9	(a)	(i)	non-	inverting (amplifier)		B1	[1]
		(ii)	(G =) 1 + R_2 / R_1		B1	[1]
	(b)	(i)	•	= 1 + 100 / 820 ut = 17 mV		C1 A1	[2]
		(ii)	(<i>R</i> ₂ / (1 +	R_1 scores 0 in (a)(ii) but possible 1 mark in each of (b R_1 / R_2) scores 0 in (a)(ii) , no mark in (b)(i) , possible 1 R_2 / R_1) or R_1 / R_2 scores 0 in (a)(ii) , (b)(i) and (b)(ii))	A1	[1]	
10	(a)	(i)	dens	sity × <u>speed of wave</u> (in the medium)		B1	[1]
		(ii)	ρ = =	(7.0 × 10 ⁶) / 4100 1700 kg m ⁻³		A1	[1]
	(b)	(i)	I = I	$T + I_R$		B1	[1]
		(ii)	1. α	$= (0.1 \times 10^{6})^{2} / (3.1 \times 10^{6})^{2}$ = 0.001		C1 A1	[2]
			2. α	≈ 1		A1	[1]
	(c)	eith or		very little transmission at an air-skin boundary (almost) complete transmission at a gel-skin boundary when wave travels in or out of the body no gel, majority reflection with gel, little reflection when wave travels in or out of the body		M1 M1 (M1) (M1) (A1)	[3]
11	(a)	(i)	unwa	anted random power / signal / energy		B1	[1]
		(ii)	loss	of (signal) power / energy		B1	[1]
	(b)	(i)	eithe	er signal-to-noise ratio at mic. = $10 \log (P_2 / P_1)$ = $10 \log (\{2.9 \times 10^{-6}\} / \{$	3 4 × 10 ⁻⁹ \)	C1	
				= 29 dB maximum length = (29 – 24) / 12 = 0.42 km = 420 m		A1 C1 A1	[4]
			or	signal-to-noise ratio at receiver = 10 lg (P_2 / P_1) at receiver, 24 = 10 lg $(P / \{3.4 \times 10^{-9}\})$ $P = 8.54 \times 10^{-7}$ W power loss in cables = 10 lg $(\{2.9 \times 10^{-6}\} / \{8.54 \times 10^{-6}\})$	0 ⁻⁷ })	(C1) (A1) (C1)	
				= 5.3 dB length = 5.3 / 12 km = 440 m		(A1)	

Pa	age 6	6 Mark Scheme: Teachers' version		Syllabus	Paper	•
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	couple	n amplifier ed to the m ater amplifi	icrophone ers scores no mark)		M1 A1	[2]
12 (a)	satellite re signal amp at a differe different fr e.g. of free	ceives great olified and ent (carrier) equencies quencies us	hitted from Earth to satellite atly attenuated signal transmitted <u>back to Earth</u>) frequency prevent swamping of uplink signal sed (6/4 GHz, 14/11 GHz, 30/20 GHz) any two other for additional physics)	(1) (1) (1) (1)	B1 B1 B2	[4]
(b)	advantage disadvanta	e.g.	because orbits are much lower		M1 A1 (M1) (A1) M1	
			more satellites required for continuous of	peration	A1	[4]